

APEX

Asteroid Probe Experiment

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Small Satellite Conference

Technical Session VI: Science Mission Payloads 1

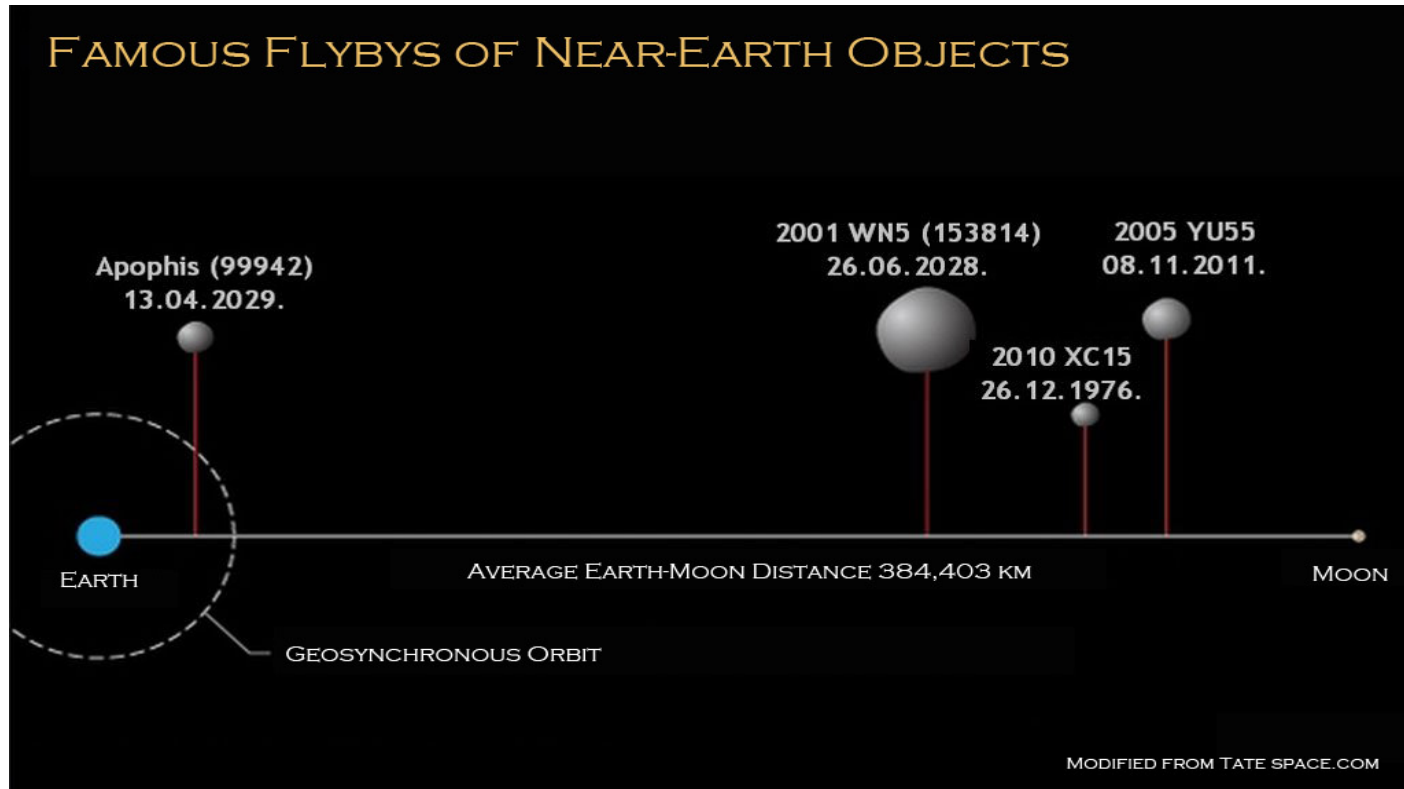
August 8, 2017



Outline

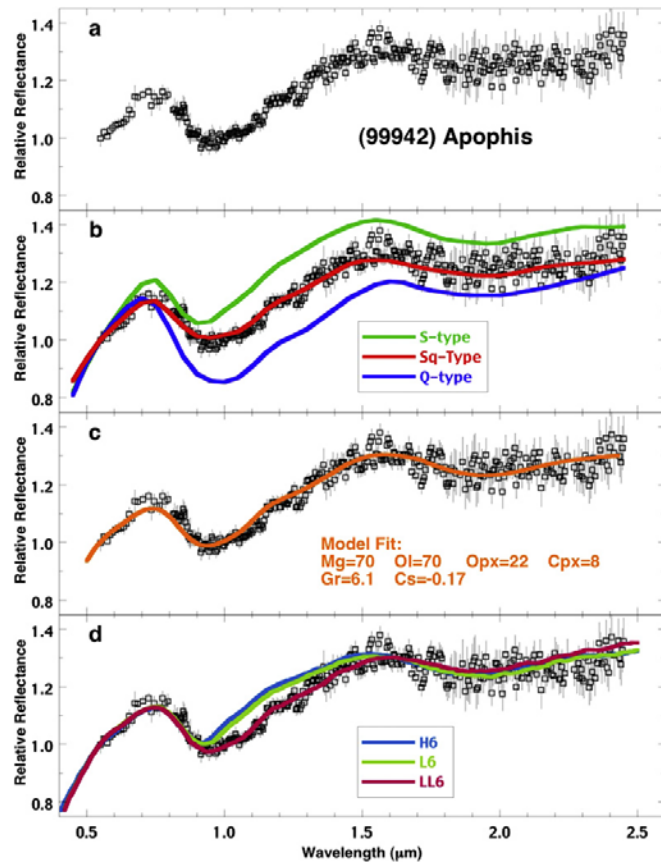
- Apophis
- Science Requirements
- Mission Requirements / Challenges
- Conops
- Spacecraft
- Summary

Apophis Encounter – Friday the 13th April 2029



Utah Sate University, Logan, UT

Apophis



65-75% olivine (70%)
17-27% orthopyroxene (22%)
3-13% clinopyroxene (8%)

Sq type asteroid

Uncommon spectrally immature inner belt asteroid with olivine and pyroxene bands

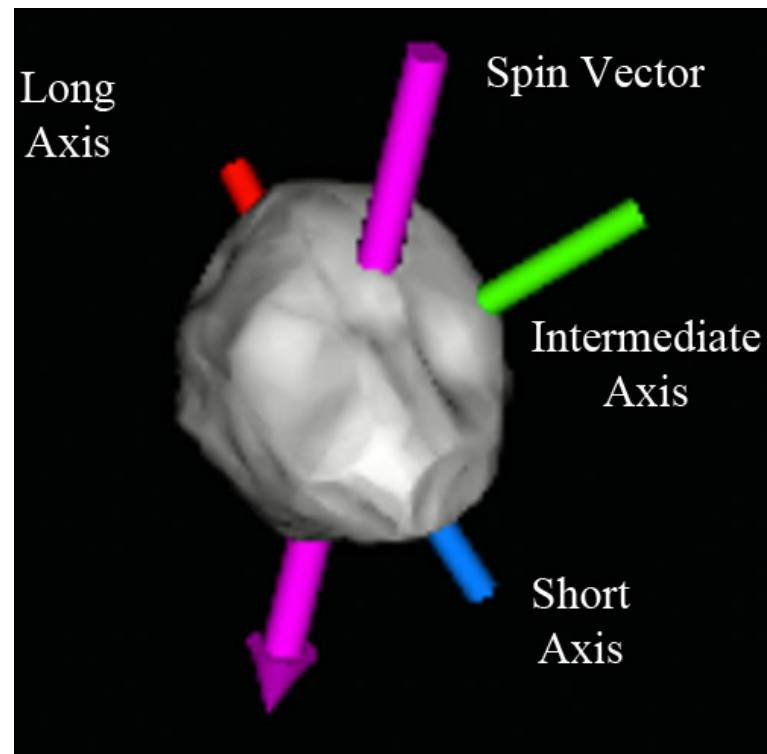
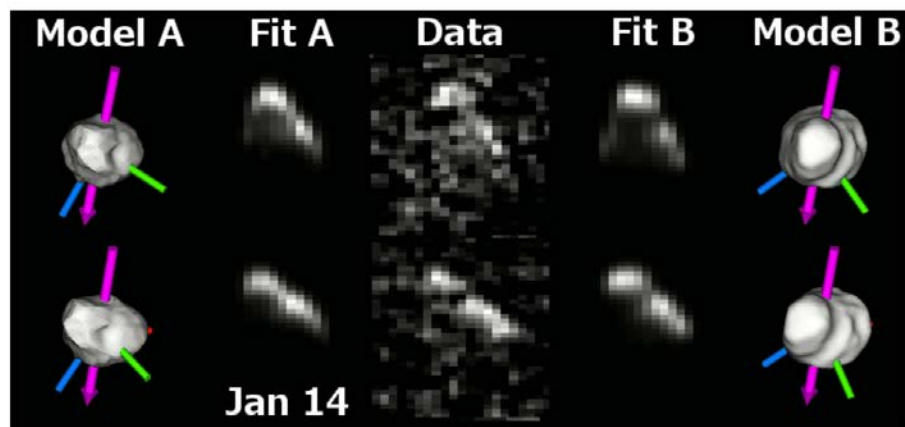
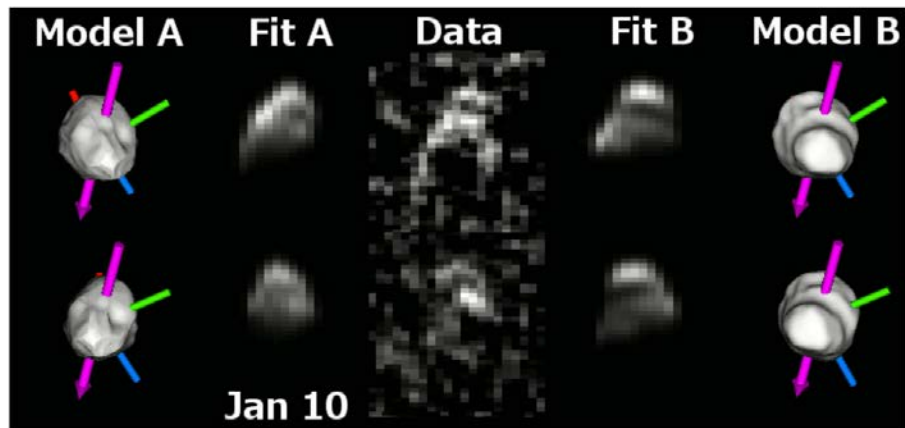
Surface possibly refreshed and rotational period and axis altered by close passage to Earth

Properties

Dimensions:	410 x 350 x 317 m
Equivalent Diameter:	340 m
Surface Area:	0.39 km ²
Volume:	0.02 km ³
Rotation:	30.4 hr (retrograde) tumbling
Precession:	262.7 days
Gravity:	0.00027 m s ⁻²

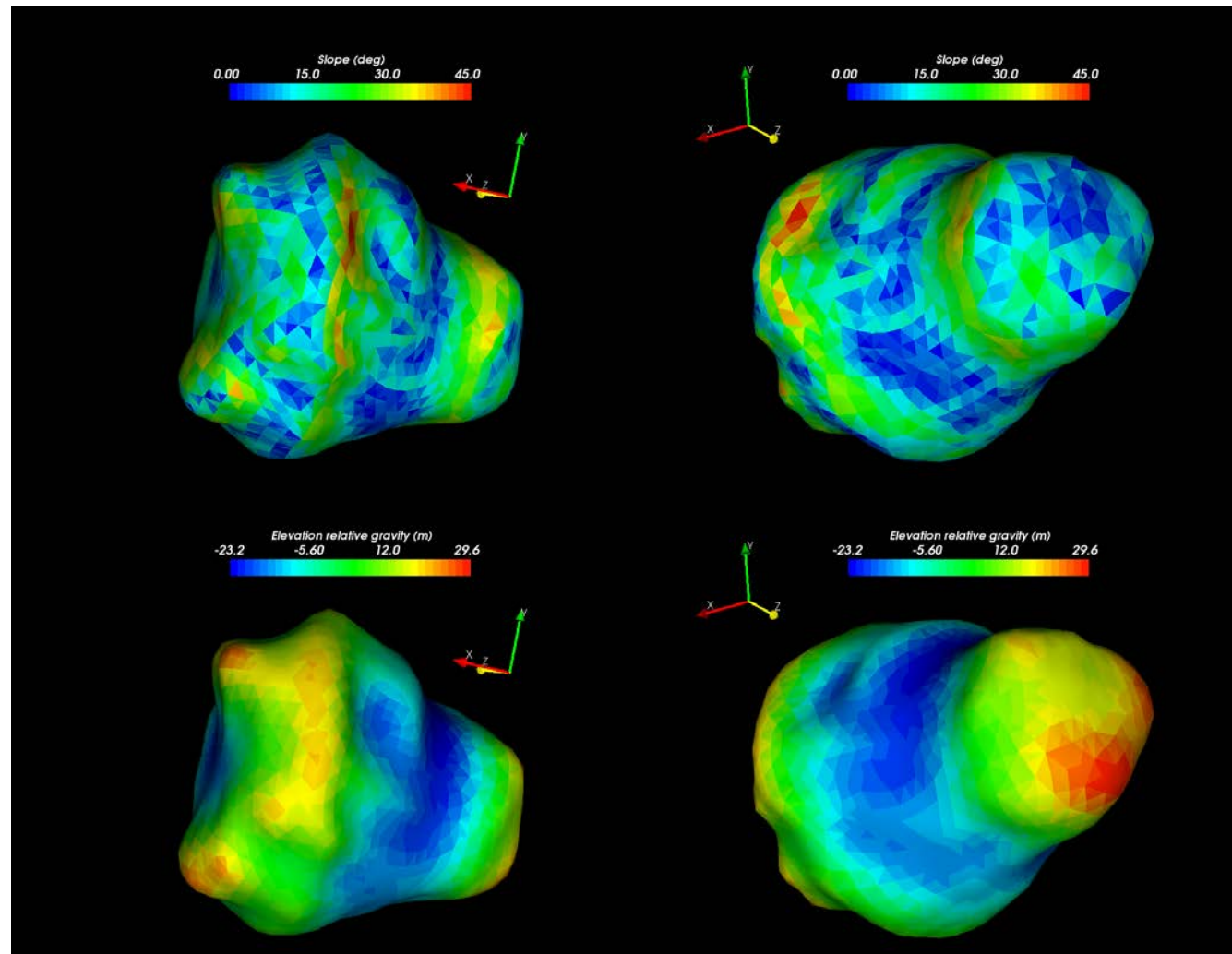
Orbital period:	324 days
Eccentricity:	0.19
Inclination:	3.33

Apophis Shape - Radar



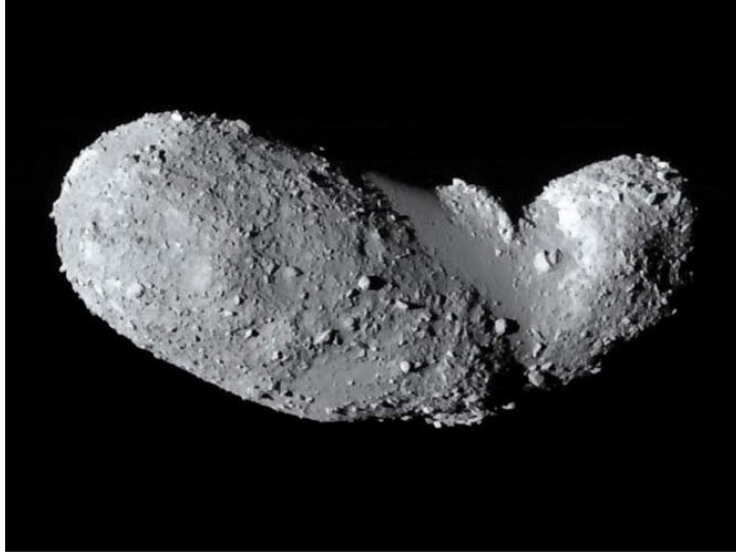
Non-Principal Axis Rotation
Period: 30.56 hr
Oscillation Long Axis: 262.7 Days
Brozović et al. 2017

Apophis Shape / Gravity

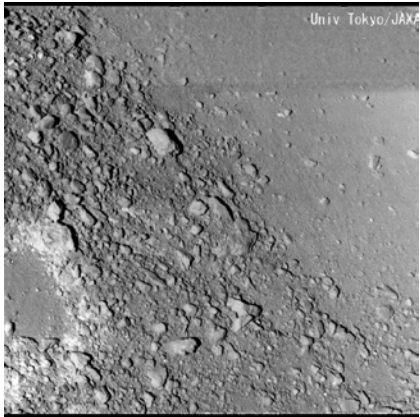


Assumes bulk density 2400 kg m^3

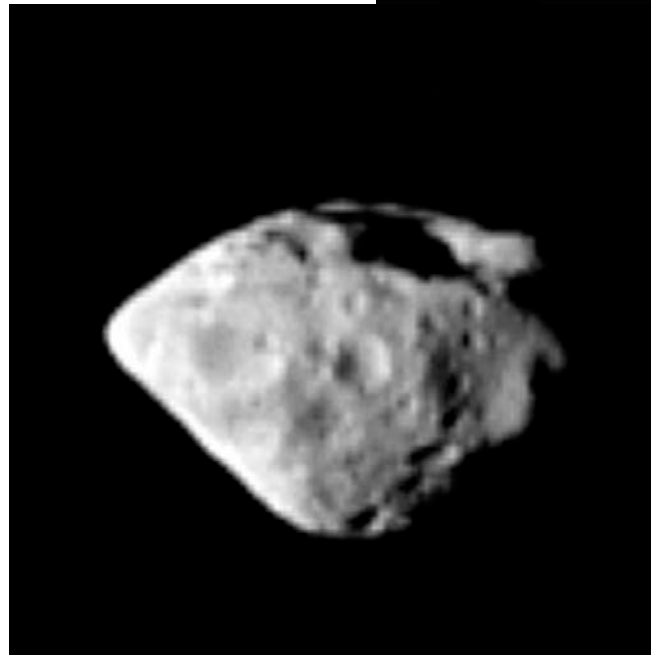
Expectations



Itokawa: 535 x 294 x 209 m



Lutetia 121000 x 101000 x 75000 m



Steins: 6670 x 5810 x 4470 m

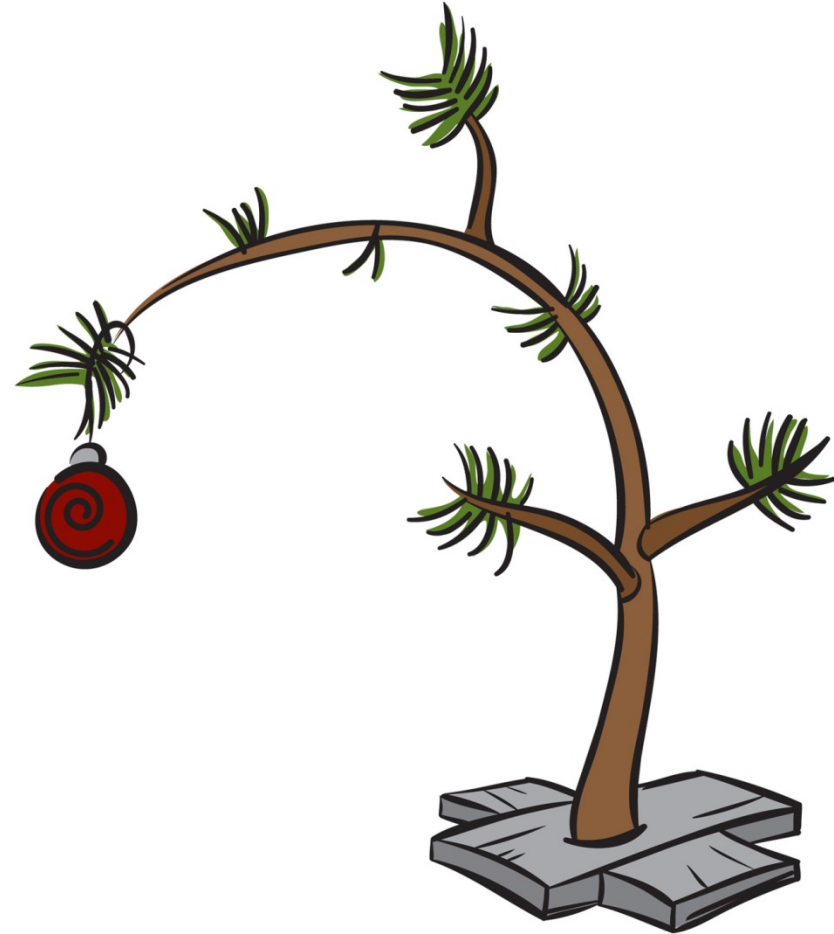


Eros: 34400 x 11200 x 11200 m

Science Requirements



Engineer / Program Manager Perspective



Scientist Perspective

Science Requirements

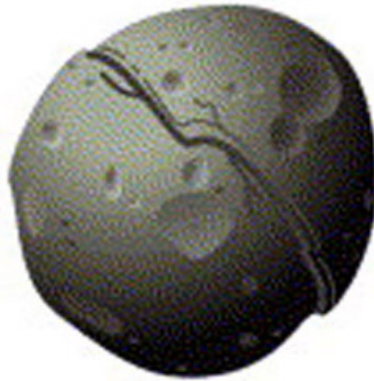
Examine the tidal distortion effects of a close encounter with the Earth in April 2029 to understand the structure – implications for history and planetary protection.

- Determine the rotational dynamics
 - Establish the physical dimensions
 - Determine the shape / topography
 - Determine the interior structure
 - Define the surface morphology
 - Define the mass
-
- Orbital change: semi-major axis 0.92 to 1.1 AU Aten to Apollo Family

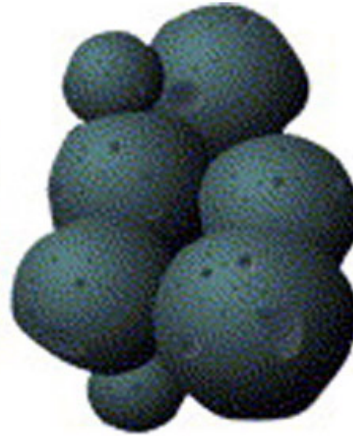
Internal Structure



Solid
Differentiated
Undifferentiated



Fractured Solid



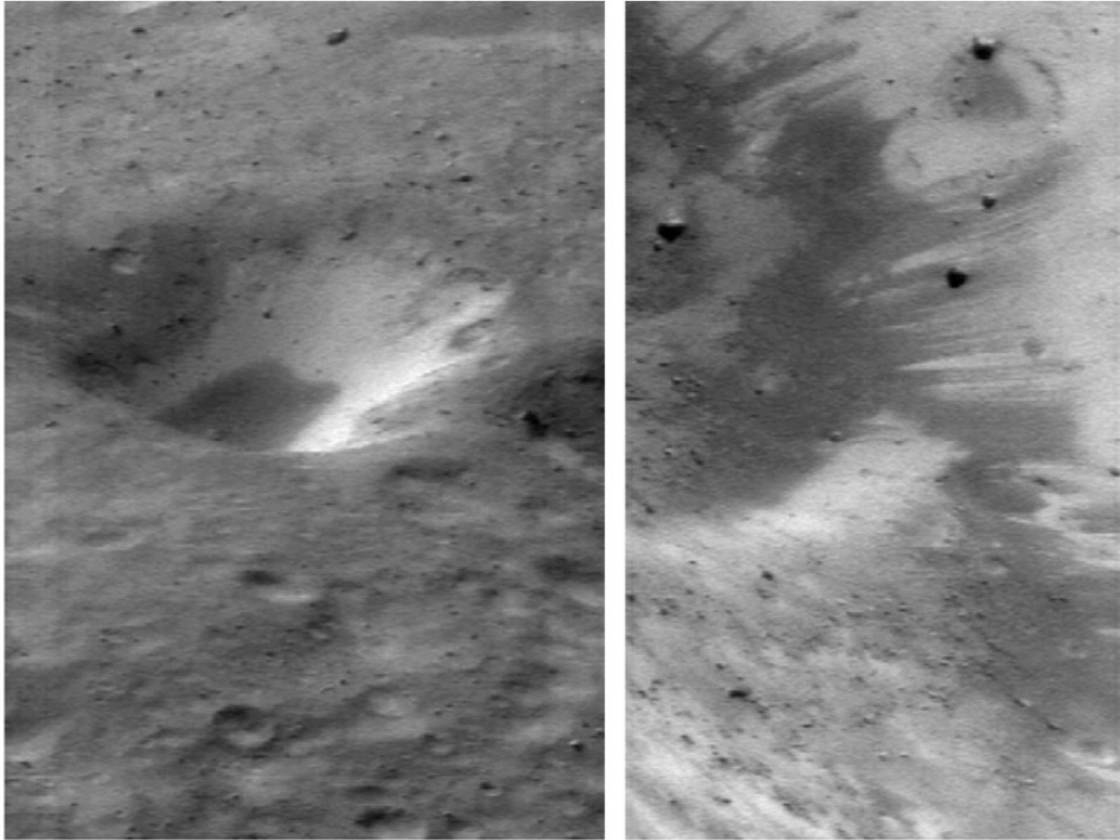
Rubble Pile
Regolith



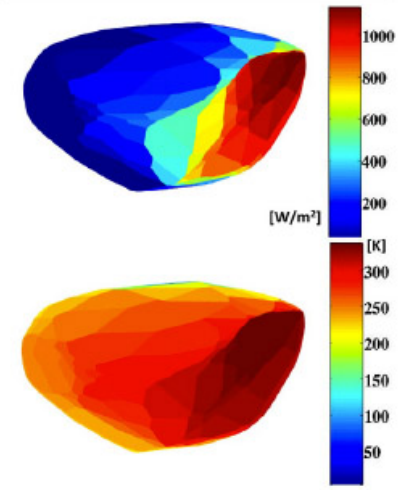
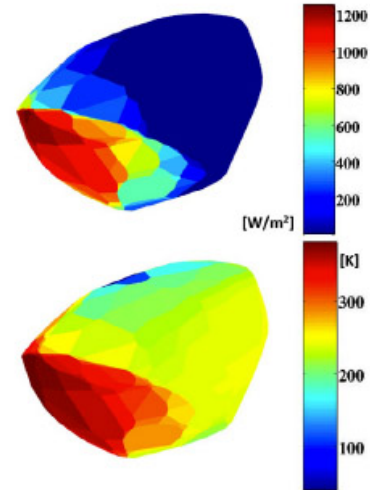
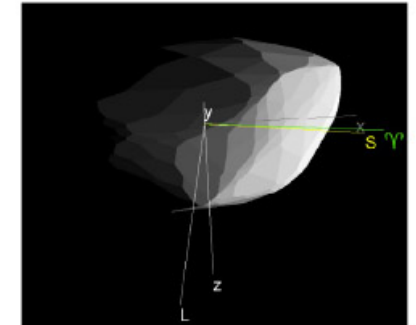
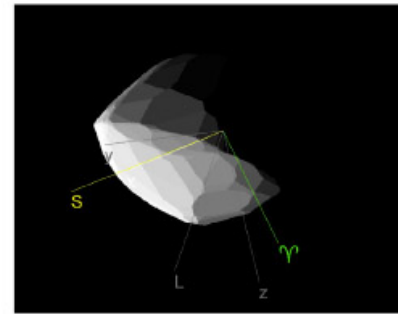
Gravel
Conglomerate

Walker et al (2006)

Surface Character and Modification



NEAR MSI 0154622520 / 0154409710



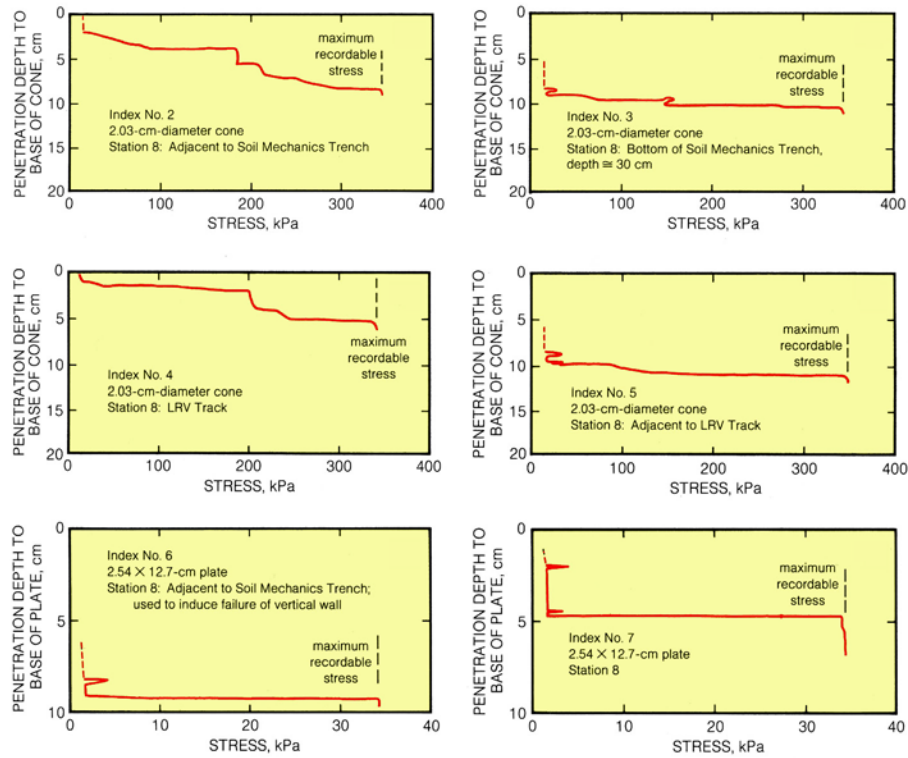
Mission Requirements / Challenges

- Rendezvous with Apophis sufficiently before Earth-encounter to map the surface and emplace a seismometer on the surface.
- Image the entire surface, in stereo, such that a DTM can be produced and the rotational dynamics established.
- Deploy seismometer on the surface to detect seismic signals induced by thermal, rotational and tidal deformation forces. Monitor baseline seismic signal and encounter signal.
- Baseline payload:
 - panchromatic imager
 - seismometer
- Additional possibilities:
 - multiband imager
 - thermal imager
 - gamma ray spectrometer

Mission Requirements / Challenges

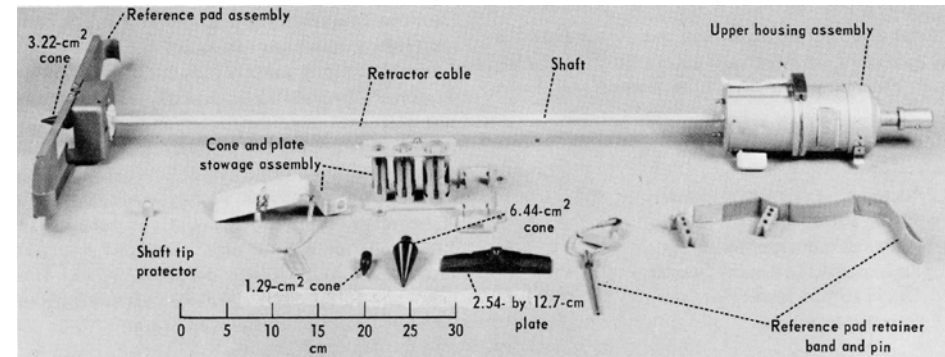
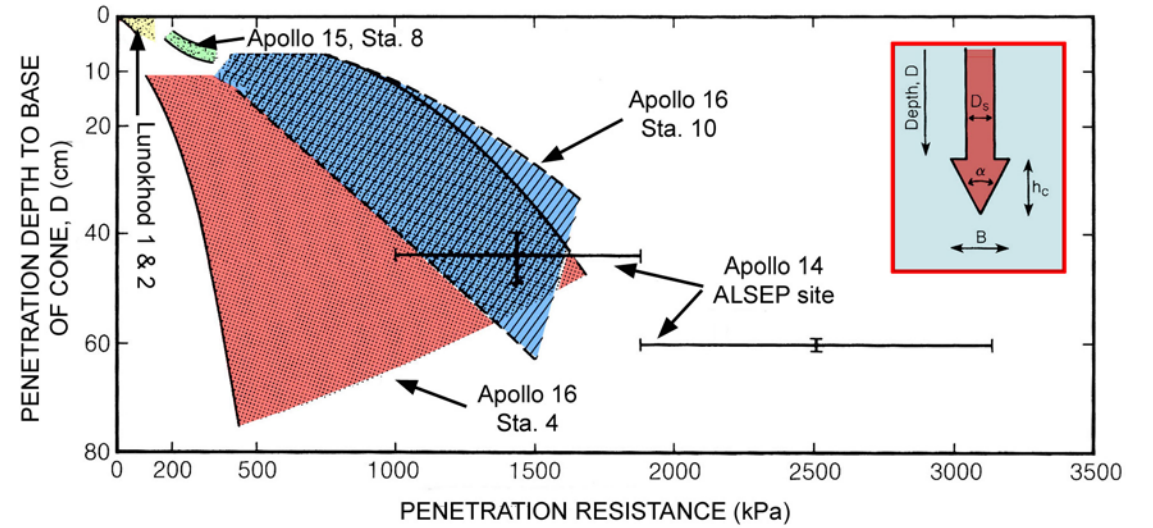
- Rendezvous and station-keep / orbit with a small body ~375 m diameter, at 1 AU, at close range (1 km). Solar forces > gravity
- CONOPS
 - “Orbit” Apophis to map the surface
 - Deploy a seismometer on the surface, ensure instrument is coupled to the surface. Touch the surface.
 - Autonomous operations
 - PROXOPS – terrain recognition
- Deployment mechanism
 - Space qualified stick - Use spacecraft to anchor the surface package
 - Launched penetrator
- Seismometer must be an independent instrument – power (solar cells and batteries), communications, electronics. Must survive for multiple months.
- Communications: Apophis to s/c, s/c to Earth
 - Data volume: Seismometer 40 Gb / Apophis day (30.4 hrs)
 - RF vs. Optical comm. for data
- PSDS spacecraft – <180 kg, low solar cell area – low power, small structure – small antenna
 - ESPA ring mount

Surface Package Anchor



Apollo 15 Cone Penetrometer Data - mare

Penetration Resistance Data



Seismometer

Molecular Electric Transducers work as a novel motion sensing mechanism by integrating mass-spring system and electro-chemical reaction.

Molecular electronic transducer

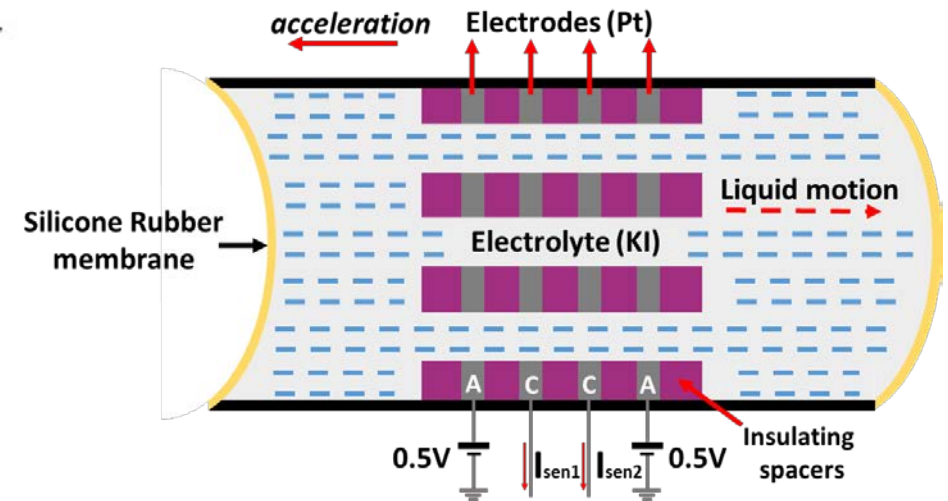
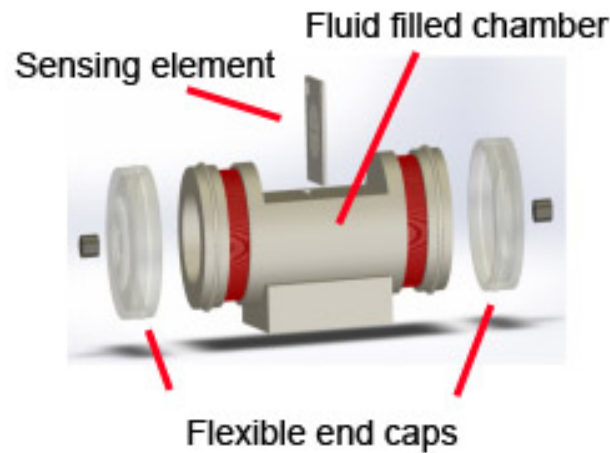
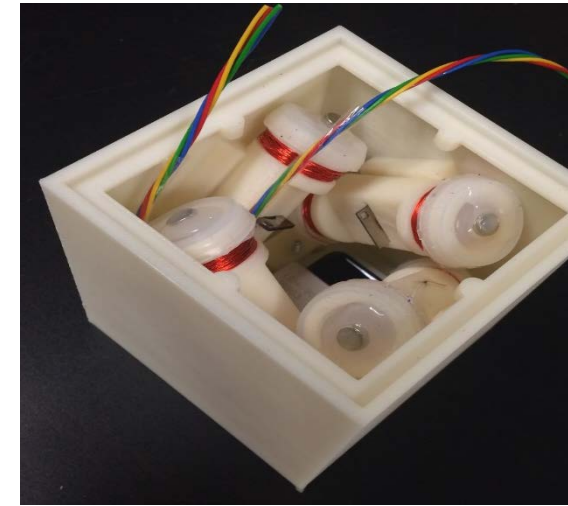
- 4 electrodes (Pt)

- Inter-electrode spacers

- Channels

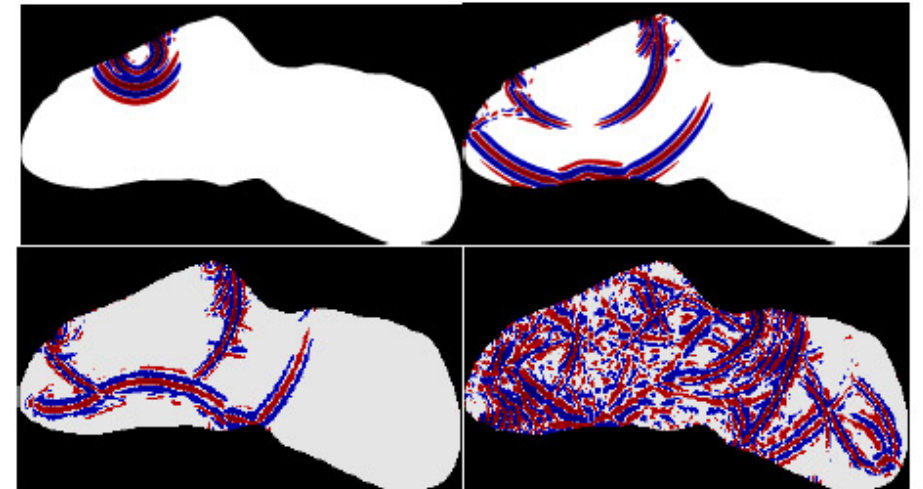
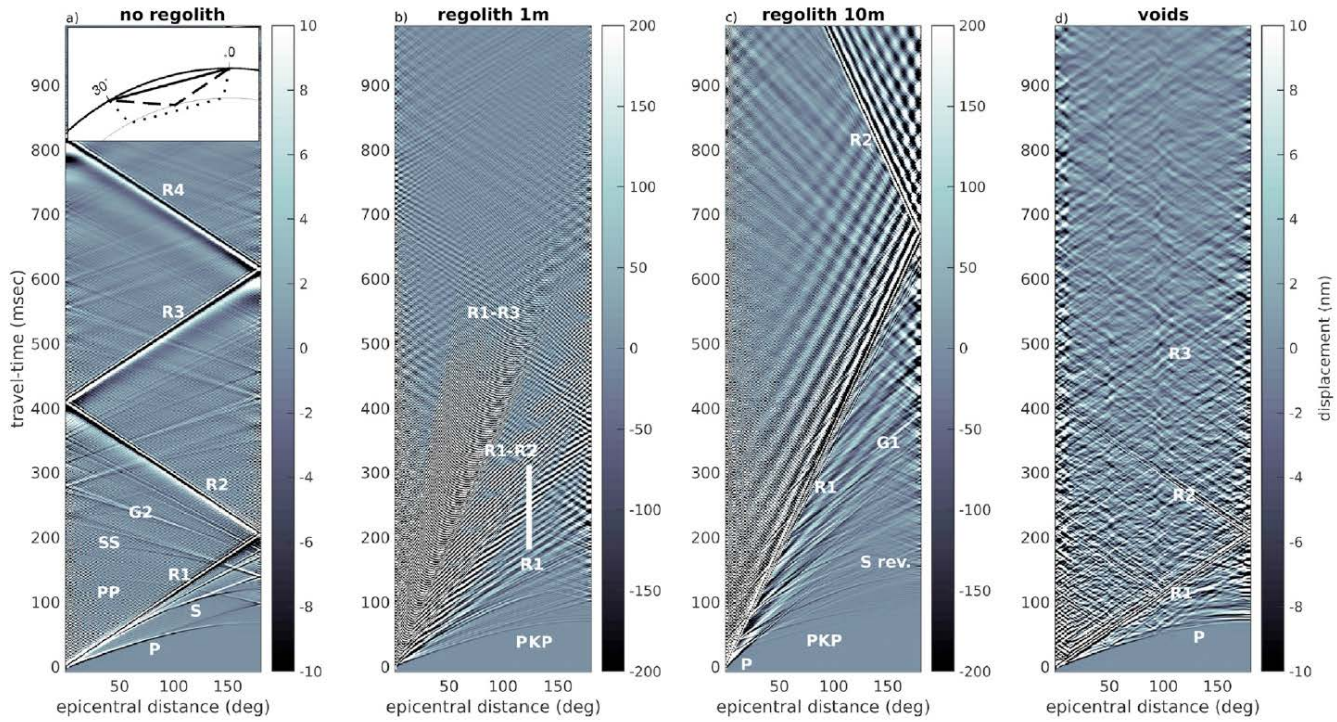
- Iodine ion solution

- Flexible diaphragm



Arizona State University

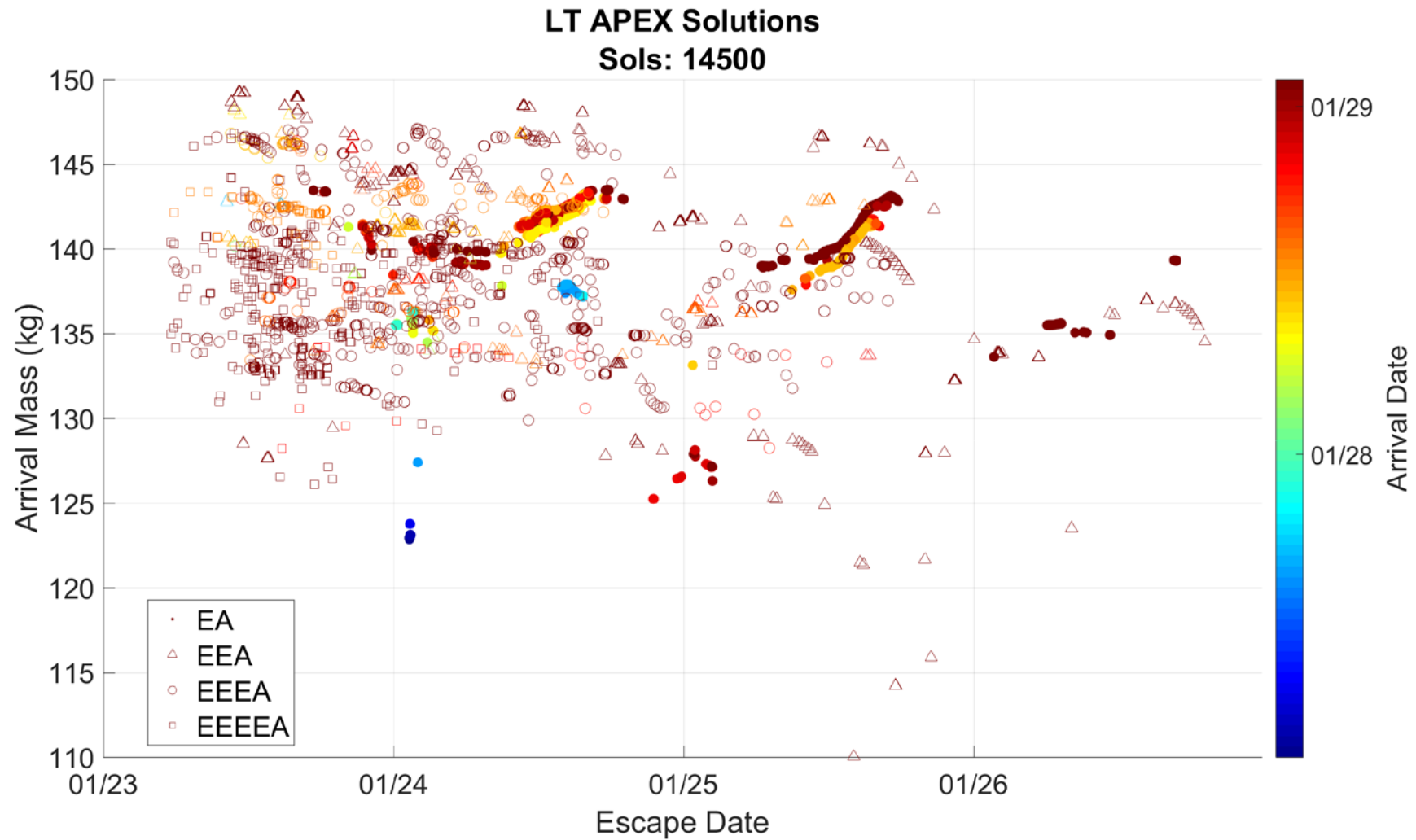
Seismology



Martin et al. (2008) 60 sec propagation

Murdoch et al. 2017

Trajectory Analysis



Trajectory Down-Select

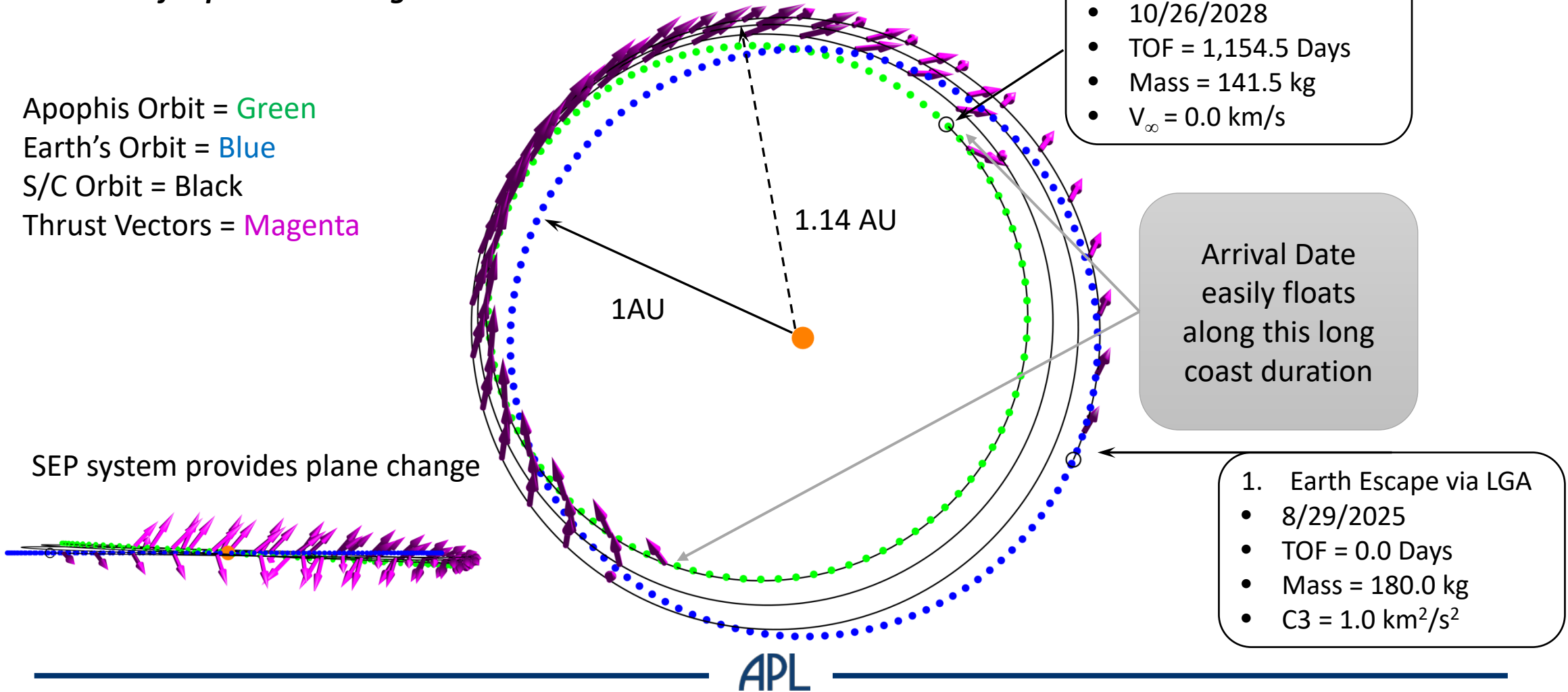
Solution	Sequence	Departure Date	Arrival Date	Time of Flight (years)	Final Mass (kg)
7105	[E]L[A]	9/3/2024	10/18/2028	4.1	143.3
8120	[E]L[A]	8/27/2025	10/23/2028	3.2	141.8
398	[E]LE[A]	6/10/2024	8/12/2028	4.2	146.8
1152	[E]LE[A]	8/26/2024	2/1/2029	4.4	148.0
670	[E]LE[A]	6/21/2025	2/1/2029	3.6	146.7
446	[E]LE[A]	7/4/2025	8/5/2028	3.1	142.9

Highlighted trajectories are mass optimal for mid-2025 launch with arrival dates several months prior to Earth close approach

Trajectory Solution 8120 [E]L[A]

Baseline for parameter engine studies

Apophis Orbit = Green
Earth's Orbit = Blue
S/C Orbit = Black
Thrust Vectors = Magenta



2. Apophis Rendezvous

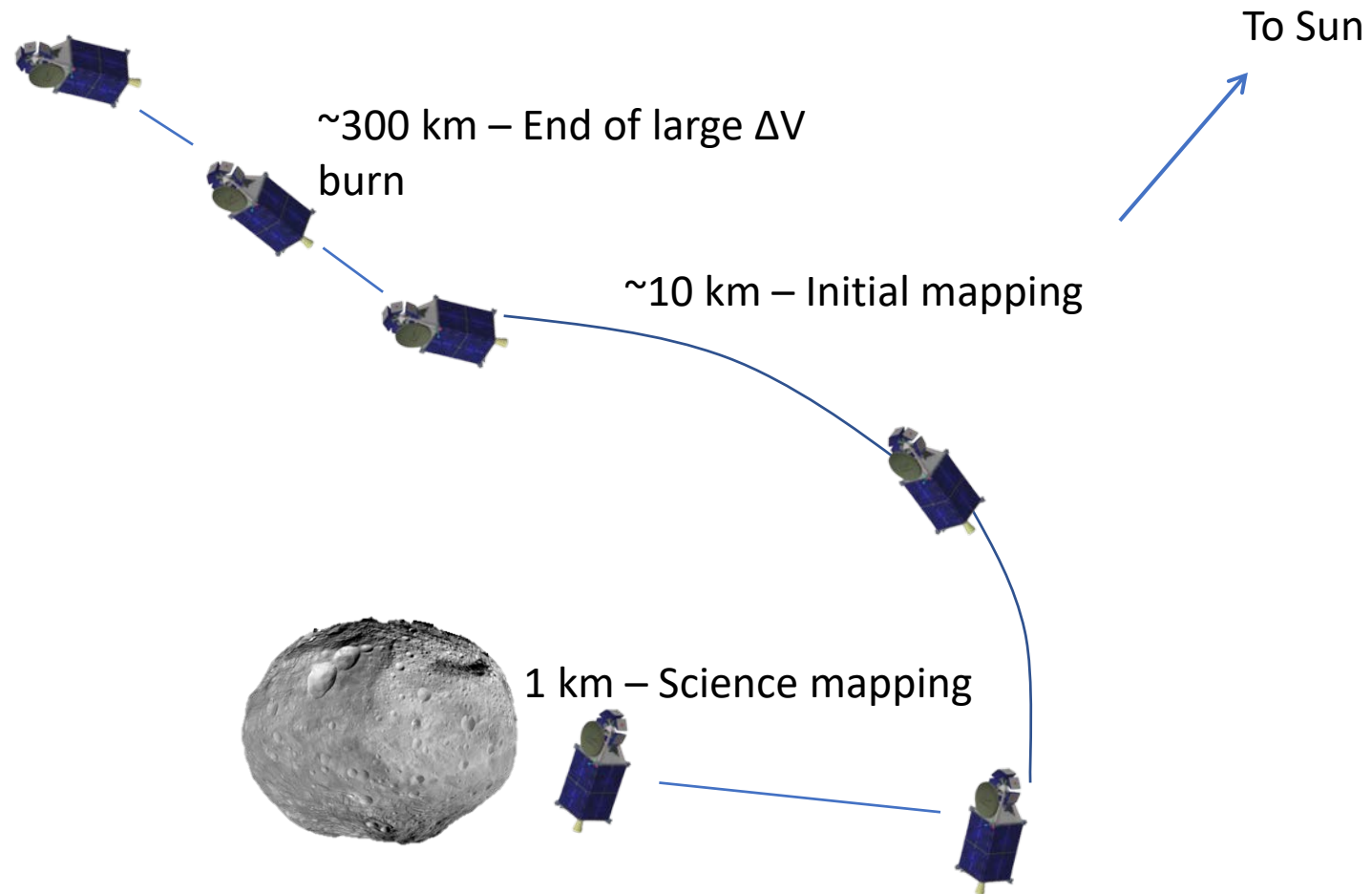
- 10/26/2028
- TOF = 1,154.5 Days
- Mass = 141.5 kg
- $V_{\infty} = 0.0$ km/s

Arrival Date easily floats along this long coast duration

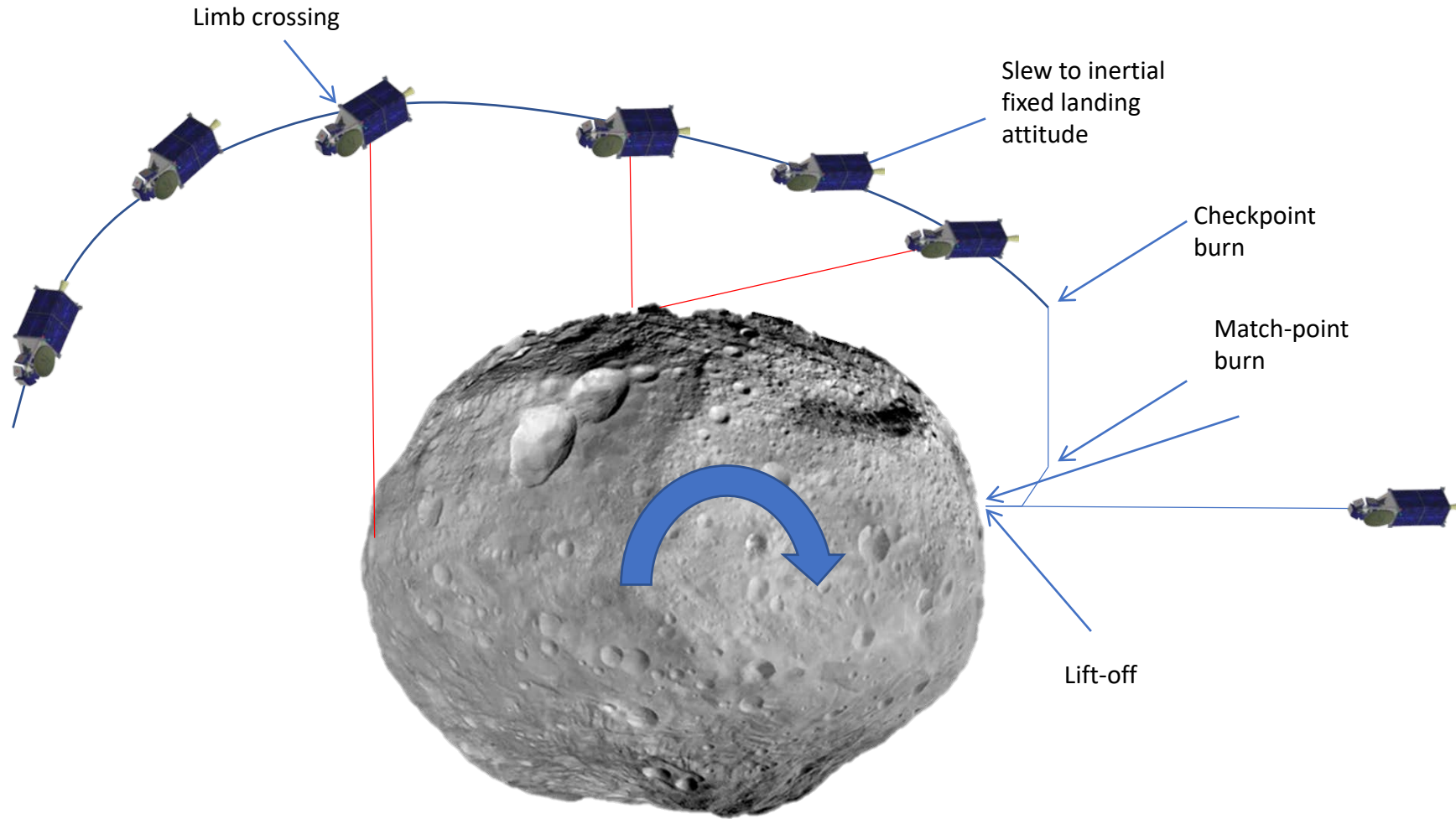
1. Earth Escape via LGA

- 8/29/2025
- TOF = 0.0 Days
- Mass = 180.0 kg
- $C3 = 1.0$ km²/s²

CONOPS - Approach and Rendezvous

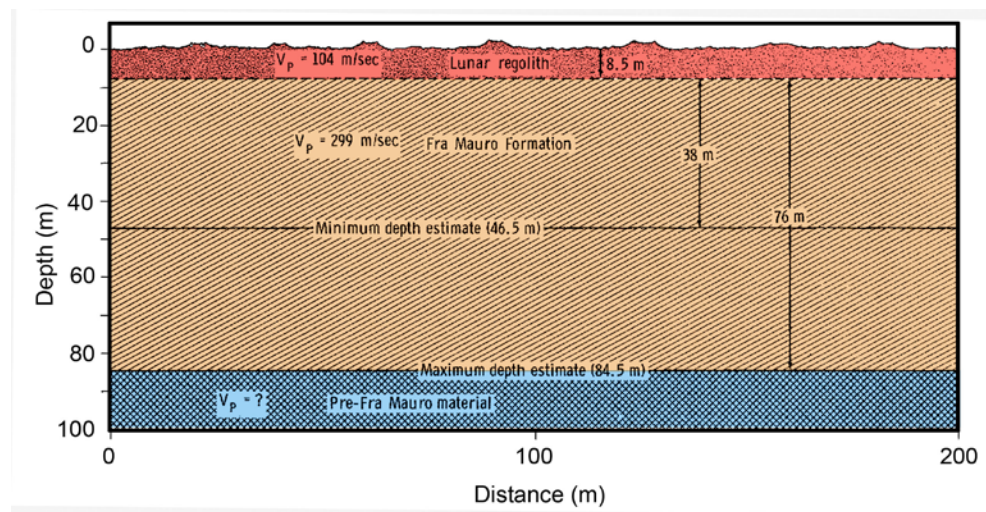
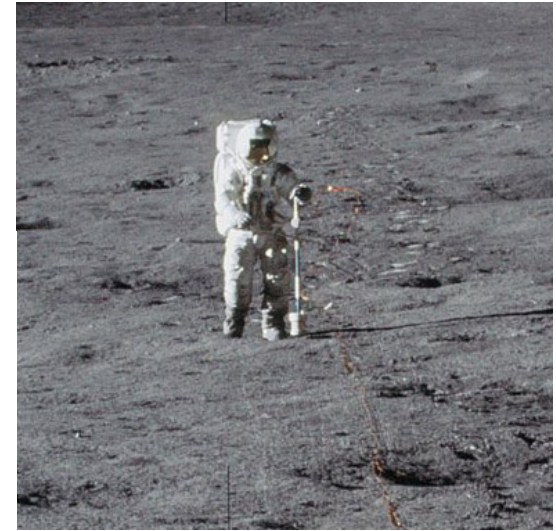
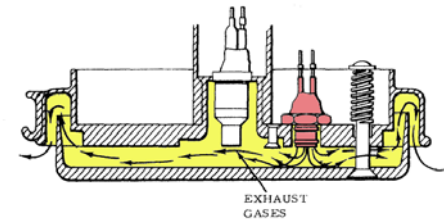
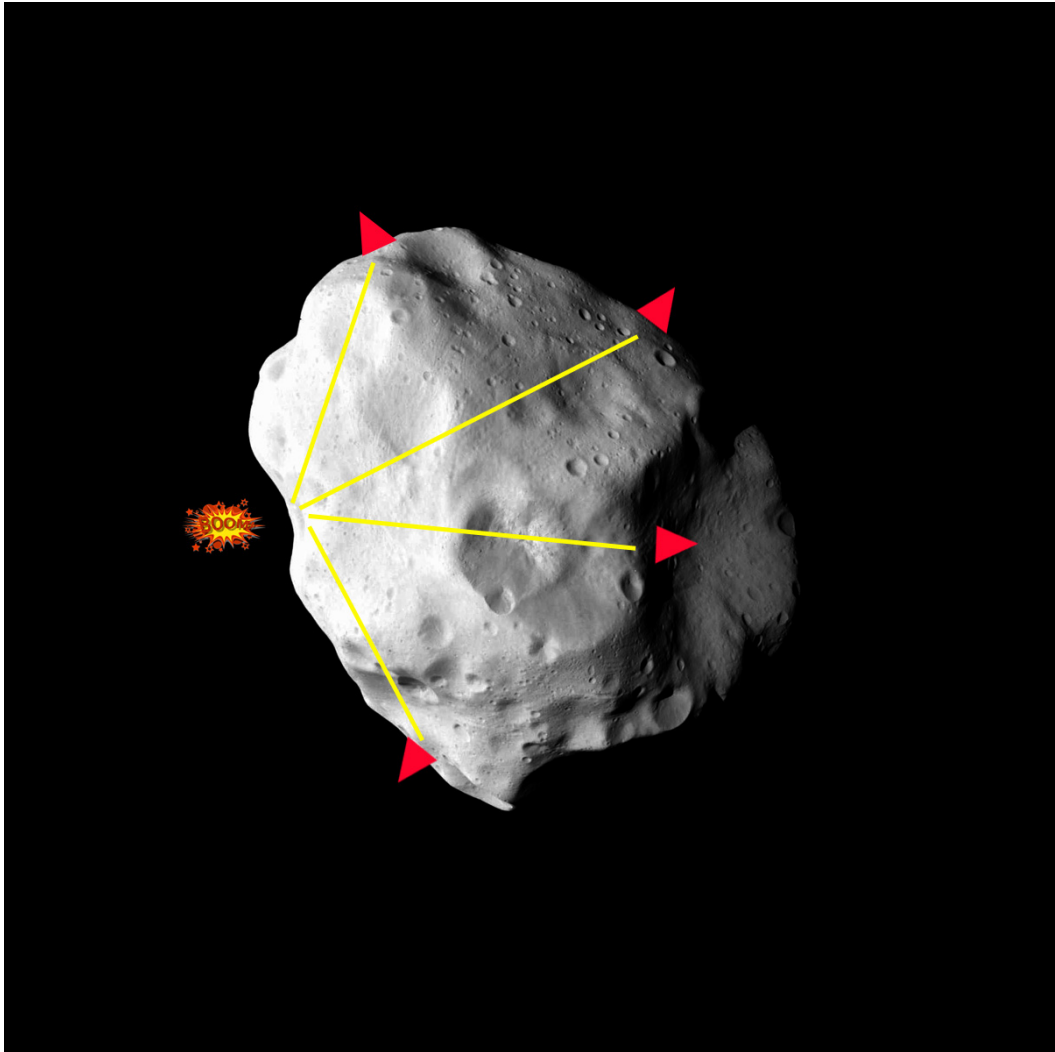


CONOPS - Surface Package Deployment



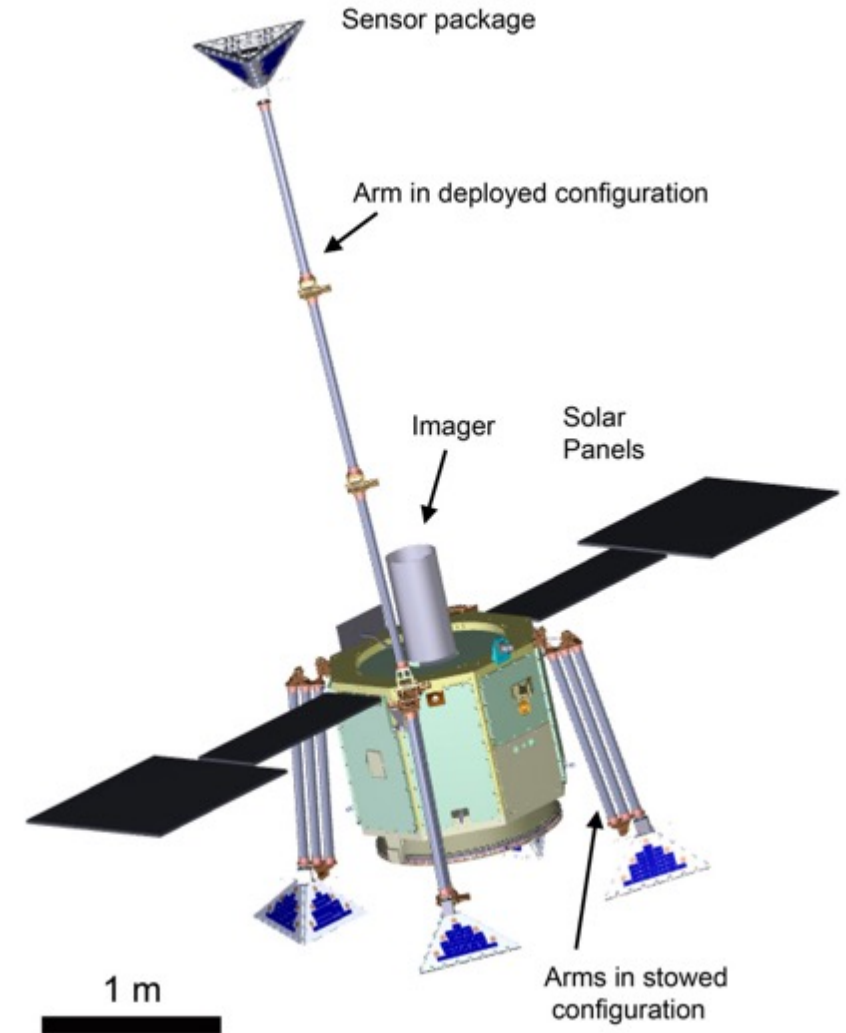
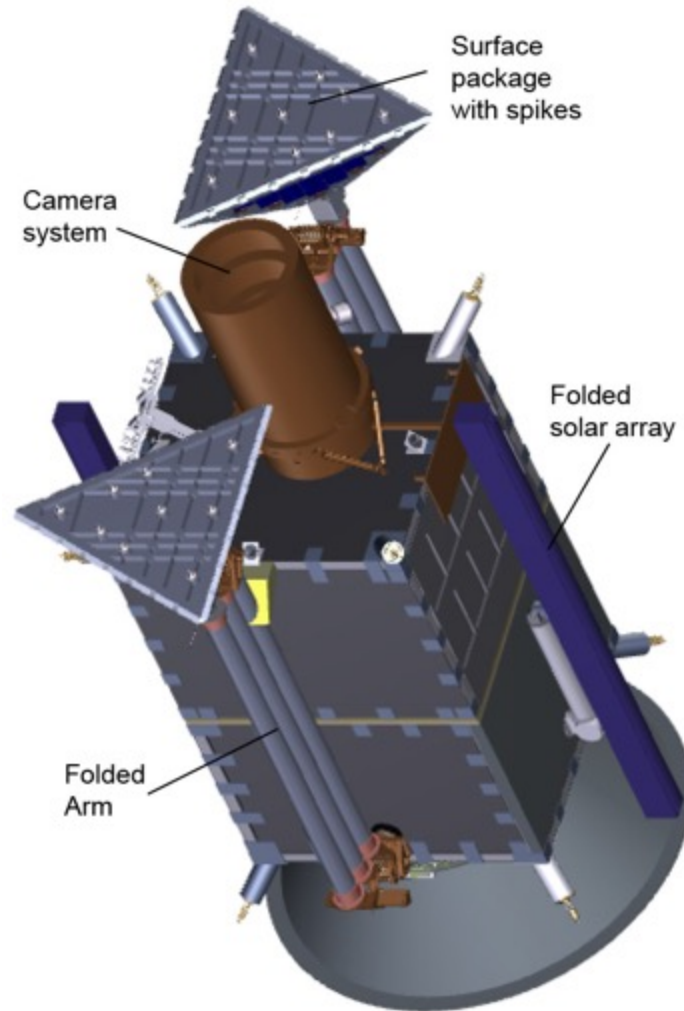
(OSIRIX-REX concept - Not to scale)

Active Seismic Experiment

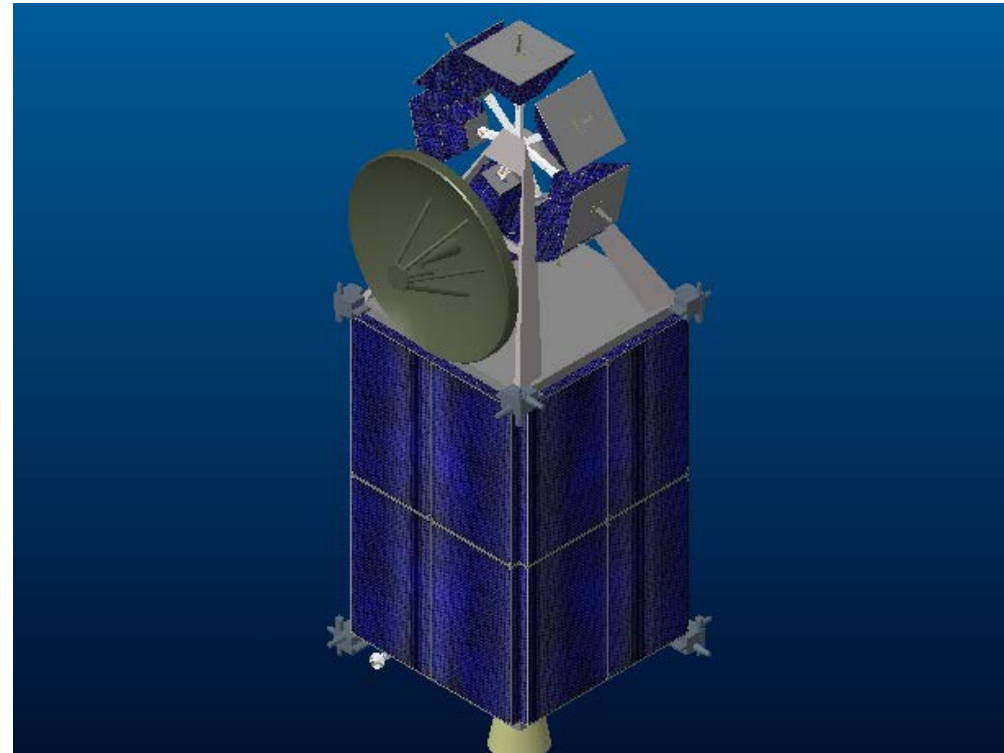
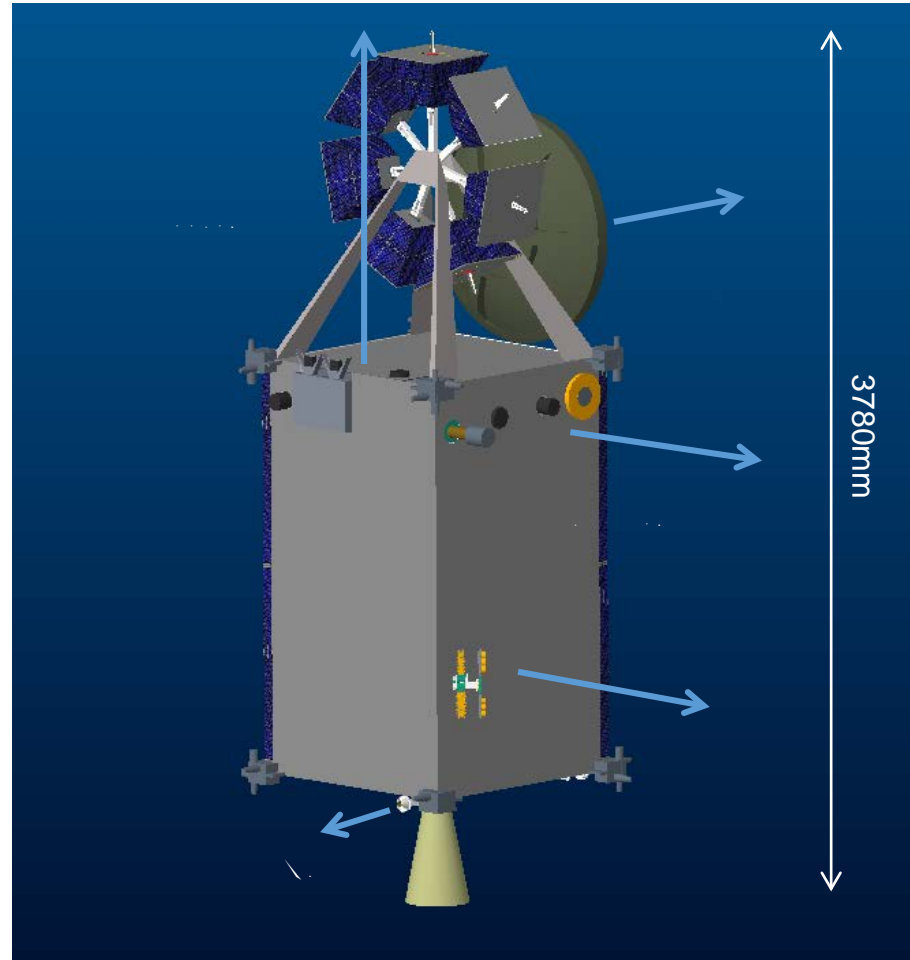


NIAC

Table 1. Flight System Mass Summary			
Subsystem	CBE (kg)	Cont. (%)	MERV (kg)
Structures	52	10	57
Integrated Propulsion	16	3	17
Avionics	11	4	11
Electrical Power	27	7	29
Attitude Determination Control	12	7	12
Thermal Control	4	15	4
RF Communications	4	6	4
Harness	11	10	12
Spacecraft Bus	136	8	146
Instrument packages	42	15	49
Total Dry Mass	174	10	195
Usable Hydrazine Propellant*			109
Propellant Residual & Pressurant			5
Total Mass*			309
Launch Capability*			358
* These values are for target 1991VG. Maximum expected resource value: MERV Contingency = Maximum expected resource value - current estimate of resource value % Contingency = [Contingency / (MERV - Contingency)] * 100			



Discovery

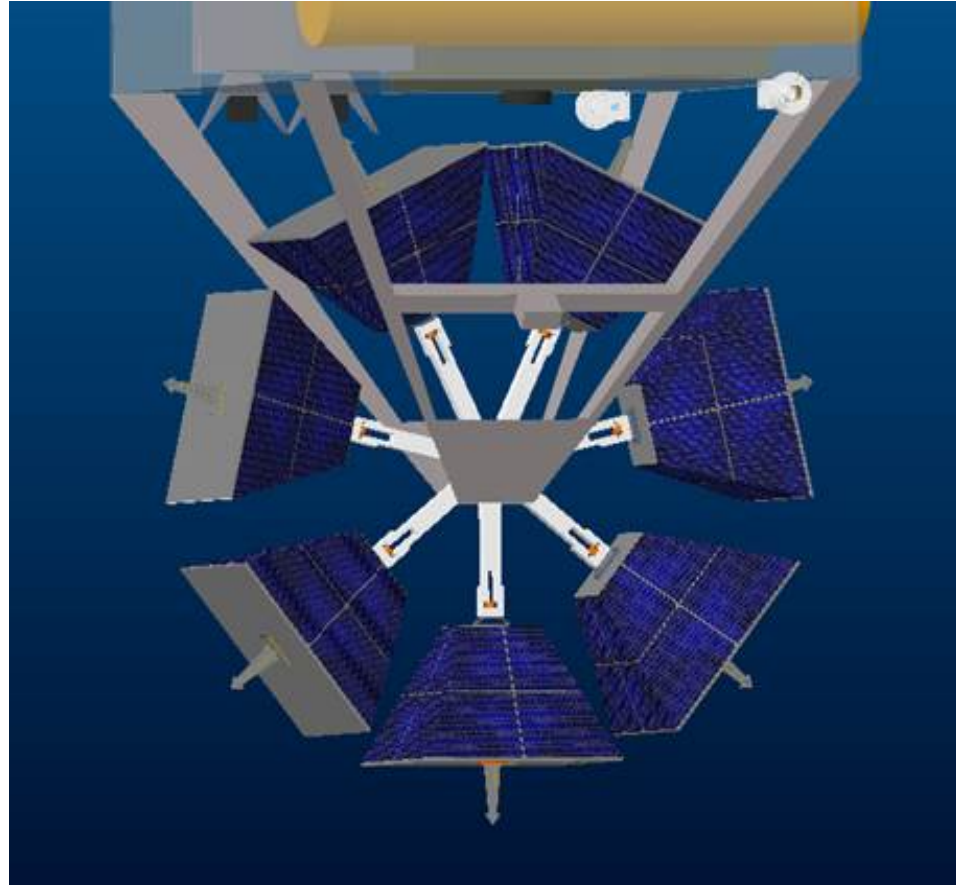


Discovery - MEL

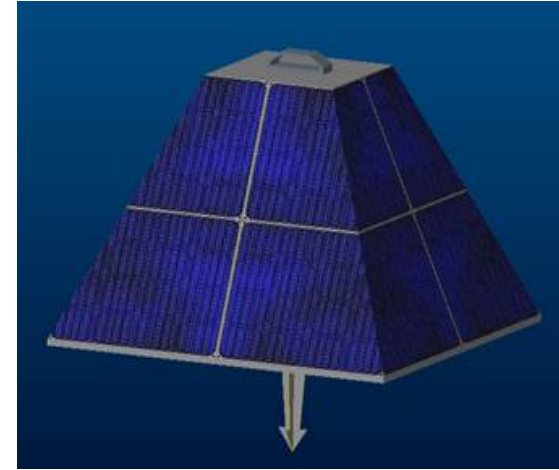
FATE Flight System Mass Summary			
Subsystem	CBE (kg)	Cont.	MEV (kg)
Structures	281.3	11%	313.5
Propulsion	56.9	6%	60.5
Avionics	28.4	6%	30.2
Electrical Power	30.3	6%	32.1
Attitude Determination and Control	57.5	3%	59.2
Thermal Control	17.9	5%	18.8
RF Communications	58.8	9%	64.1
Harness	42.4	3%	43.6
Spacecraft Bus Total	573 kg	8%	622 kg
Wide Angle Camera	5.5	15%	6.3
Narrow Angle Camera	8.0	15%	9.2
Seismometer	39.6	4%	41.3
Laser Altimeter	21.4	15%	24.6
Payload	75 kg	9%	81 kg
Total Dry Mass	648 kg	9%	704 kg
Maximum Possible Dry Mass (used to size propulsion)			850 kg
		Consumables	513 kg
Apophis Observatory Total Wet Mass			1360.0 kg

Launch Vehicle Capability
 (C3 of -0.7 km²/s²; Antares 232)
 1805kg

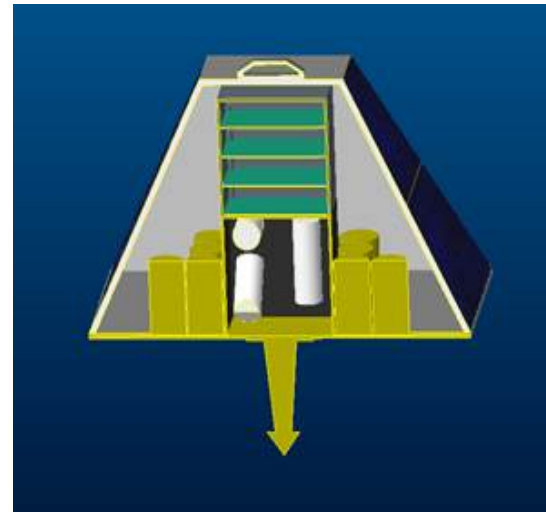
Discovery - Deployment



Ferris Wheel – Sensors and Sources



Surface package



Summary - APEX

- Challenging mission
- PSDS – low mass, small budget
- CONOPS at a small no-g body – Autonomous ops
 - Rendezvous
 - Touch the surface
- Emplacement of seismometer (tiny self-contained “spacecraft”) – ensure coupling
- Data downlink – high data volume from seismometer
- Close to the Earth, close to the sun
- Looking for suggestions

- *Acknowledgements:* NASA Planetary Science Deep Space SmallSat Studies Program, NASA Innovative Advanced Concepts Program, Applied Physics Laboratory, Cast of Summer Interns